CATION DIFFUSION IN CARBONATE MINERALS: DETERMINING CLOSURE TEMPERATURES AND THE THERMAL HISTORY FOR THE ALH 84001 METEORITE. D. K. Fisler, R. T. Cygan, and H. R. Westrich, Geochemistry Department, Sandia National Laboratories, Albuary NM 87185-0750

The diffusion rates of cations in carbonate minerals can be used to help constrain the thermal conditions expected for the formation and subsequent history of the ALH 84001 Martian meteorite. Since homogenization of the carbonate globules readily occurs at elevated temperatures—the extent being a function of grain size and diffusion coefficient for a given temperature-one can examine the measured compositional zoning patterns of the carbonate phases to see how the profiles might be preserved or modified [1-6]. Of course, thermal cooling histories of rocks can be quite complex and will need appropriate diffusion models in order to correctly interpret the compositional zoning patterns [7]. A convenient method based on the concept of closure temperature is prvided by Dodson [8]. He derived an approximation for the relationship between cooling rate, grain size, and closure temperature for thermally-activated diffsion in geological systems. The diffusion-based cooling model is given by the following nonlinear equation:

$$\frac{\mathrm{E}_{a}}{\mathrm{R}\mathrm{T}_{c}} = \ln\left(\frac{-\mathrm{A}\mathrm{R}\mathrm{T}_{c}^{2}\mathrm{D}_{o}}{\mathrm{a}^{2}\mathrm{E}_{a}\mathrm{s}}\right)$$

where \mathbf{E} is the activation energy for diffusion, R is the gas constant, Ts the closure temperature, D the diffusion coefficient, and A is a geometric factor (equal to 55 for a sphere), and s is the cooling rate. This equation provides a series of closure tempeatures as a function of cooling rate and grain The closure temperature represents that radius. temperature below which a cation zoning pattern would be essnetially frozen or preserved, as observed in the carboate globules found in the Martian meteorite. The approximation includes the assumption that the stym under consideration cools to a temperature low enough that diffusion is no longer significant over the life of the system.

Unfortunately, few data for the diffusion rates of cations in carbonate minerals are available that can be directly used to determine closure temperatures. However, our experimental approach that uses thin film-mineral diffusion couples [9] provides an accurate means of obtaining relatively low temperature diffusion data for cations. We have used a high-vacuum evaporator to deposit a thin film (~1000 Å) coating of an isotoqually-enriched tracer onto a cleaved (104) calcite surface. The calcite

The interface between the enriched coating and the mineral suface is identified by a discontinuity in the depth profile. The raw concentration profile is then transformed to concentratioersus depth using a sputtering rate (2.85 Å/sec) determined by measuring the depth of a sputtered crater using a profilometer on an uncoated sample. The diffusion coefficient is obtained by fitting the final diffusion profile to an error function solution to Fick's diffusion equations, æbendent on surface concentration C bulk concentation C_b, and the diffusion coefficient D [10]. The variation of concentration as a function of depth x and experimental time t is described by:

$$C_x = C_s + (C_b - C_s) \operatorname{erf}\left(\frac{x}{4\sqrt{Dt}}\right)$$



Figure 1. Raw ion microprobe profile as function of sputtering time $for^{44}Ca$ in calcite annealed at $6\mathfrak{G}0$ for 510 hours. The vertical line shows the location of the interface between the enriched coating and the calcite interface.



Figure 2. Diffusion profile for ${}^{44}Ca$ in calcite annealed at 65°C for 510 hours. The best fit (solid line) of the diffusion model to the observed data is based on the optimization of the_b, CC_s, and D parameters.

Figure 2 provides an example of the calcium diffusion profile from one of our preliminary experiments performed at 6500 for 510 hours. The profile exhibits the expected decrease of the calcium tracer (⁴⁴Ca) content up to a depth opproximately 0.15 μ m. The best fit of the observed profile to the diffusion model given above is also shown in the figure, and provides a calcium self-diffusion coefficient of 1.1 x 10^{21} m²/sec.

Experiments were performed for calcium and magnesium diffusion in calcite at tempteures from 500° - 700°C. Values for the activition energy F and pre-exponential term D for the Arrhenius relation were obtained, and then used to calculate the expected closure temperatures for various grain sizes and cooling rates expected for the carbonate globules in the ALH 84001 meteorite.

Using the data for calcium diffusion in calcite [11], cooling rates must be more rapid than 10000 for the cation zoning to have been preserved if the Martian carbonates formed at more than^o650Our preliminary experimental data indicate that magnesium diffusion rates in calcite (D = $10m^2/s$ at 700°C) are more than two orders-of-magnitude faster than those for calcium. The closure temperatures based on the magnesium diffusion data are lower and therefore require a more rapid cooling rate than those calculated based on the calcium diffon data (see Figure 3). This analysis demonstrates that any model for the high temperature



Figure 3. Closure temperatures as a function of grain diameter and cooling rate calculated for calcium and magnesium diffusion in calcite. The vertical line represents the mean diameter of the carbonate globules in the ALH 84001 meteorite.

formation of the carbonate assemblage requires rapid cooling for the preservation of the observed chemical zoning. However, a low temperature model for formation is not constrained by the experimental diffusion data.

References: [1] Harvey, R. and McSween, H. Y. (1996) Nature, 382, 49-51. [2] Mittle feldt, D. W. (1994), Meteoritics, 29, 214-221. [3] Treiman, A. H. (1995), Meteoritics, 30, 294-302. [4] Thomas-Keprta, K. L. et al. (1997), LPSC, in press. [5] McKay, D. S., et al. (1996). Science, 273, 924-930. [6] Romanek, C. Set al. (1994), Nature, 372, 655-657. [7] Ganguly, J. and Tazzoli, V. (199 Geoch. Cosmoch. Acta, 58, 2711-2723. [8] Dodson, M. H. (1973), Contrib. Min. Pet., 40, 259-274. [9] Schwandt, C. Set al. (1993), Pure Applied Geoph., [10] Crank, J. (1975)The 103, 631-642. Mathematics of Diffusion, Oxford University Press, London. [11] Farver, J. R. and Yund, R. A. (1996) Contrib. Min. Pet., 123, 77-91.

Acknowledgments: We are grateful for the technical discussions with Kathie Thomas-Keprta and Craig Schwandt that led to the application of our experimental diffusion data to the Martian meteorite. This research was supported by the U. S. Department of Energy, Office of Basic Energy Sciences, Geosciences Research, under contract DE-AC04-94AL85000 with Sandia National Laboratories.